## **RESEARCH**



# Toward a globally lunar calendar: a machine learning-driven approach for crescent moon visibility prediction

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## **Abstract**

This paper presents a comprehensive approach to harmonizing lunar calendars across diferent global regions, addressing the long-standing challenge of variations in new crescent Moon sightings that mark the beginning of lunar months. We propose a machine learning (ML)-based framework to predict the visibility of the new crescent Moon, representing a signifcant advancement toward a globally unifed lunar calendar. Our study utilized a dataset covering various countries globally, making it the frst to analyze all 12 lunar months over a span of 13 years. We applied a wide array of ML algorithms and techniques. These techniques included feature selection, hyperparameter tuning, ensemble learning, and region-based clustering, all aimed at maximizing the model's performance. The overall results reveal that the gradient boosting (GB) model surpasses all other models, achieving the highest F1 score of 0.882469 and an area under the curve (AUC) of 0.901009. However, with selected features identifed through the ANOVA F-test and optimized parameters, the Extra Trees model exhibited the best performance with an F1 score of 0.887872, and an AUC of 0.906242. We expanded our analysis to explore ensemble models, aiming to understand how a combination of models might boost predictive accuracy. The Ensemble Model exhibited a slight improvement, with an F1 score of 0.888058 and an AUC of 0.907482. Additionally, the geographical segmentation of the dataset enhanced predictive performance in certain areas, such as Africa and Asia. In conclusion, ML techniques can provide efficient and reliable tool for predicting the new crescent Moon visibility that would support the decisions of marking the beginning of new lunar months.

**Keywords:** Crescent visibility, Machine learning, Lunar calendar, Prediction, Hyperparameter tuning, Feature selection, Ensemble learning, Astronomy

## **Introduction**

Approximately 24.1% of the global population, including Muslims [\[1](#page-21-0)], adhere to the lunar calendar, which plays a crucial role in defning the start of each month, initiated by the sighting of the new crescent Moon. This method of timekeeping is integral to various religious practices around the world, underscoring the signifcance of the crescent Moon as a symbol marking religious events. Precise observation of the crescent Moon is



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pivotal, as it signifes the onset of new periods in the lunar calendar, infuencing numerous religious celebrations.

The importance of the crescent Moon in the lunar calendar can be summarized as follows; determination of months as the lunar calendar is based on the cycles of the Moon, and the sighting of the crescent Moon is used to establish the beginning of each month. The tradition of sighting the crescent Moon is a visual confirmation of the new month. It involves observing the thin crescent Moon shortly after Sunset, and the testimony of reliable witnesses is often used to confirm the sighting. The crescent Moon sighting involves the community, fostering a sense of unity and shared religious observance. In many Muslim-majority countries, committees or religious authorities are tasked with officially announcing the sighting of the crescent Moon. The sighting of the crescent Moon is especially significant during the month of Ramadan. The first sighting of the crescent Moon determines the start of Ramadan, the month of fasting, refection, and increased devotion for Muslims around the world. The sighting of the crescent Moon also plays a crucial role in determining the end of Ramadan and the celebration of Eid al-Fitr. The month of fasting concludes with the sighting of the crescent Moon, and the next day is celebrated as a day of festivity and thanksgiving. Similarly, the crescent Moon is observed to determine the beginning of the month of Dhu al-Hijjah, during which the Hajj pilgrimage takes place. The 10th day of Dhu al-Hijjah, known as Eid al-Adha, also relies on the sighting of the crescent Moon to mark the end of Hajj and the beginning of the festival of sacrifce. In summary, the crescent Moon holds signifcant importance in Islam as it is used to determine the start of each lunar month. The practice of crescent Moon sighting is deeply rooted in Islamic tradition and serves as a communal and visual way to mark the passage of time and the beginning of important religious observances.

This paper investigates the use of supervised ML techniques in predicting the visibility of the new crescent Moon. For those ML systems, there are several input features related to geographical aspects such as the longitude and the latitude, and other features such as the atmosphere, Moonset, Sun\_Moon\_lag, age\_of\_Moon, Moon\_altitude, Sun\_altitude, Moon\_azimuth, Sun\_azimuth, elongation, illumination, and conjunction time. The target or the output is the fact that the crescent was seen by the eye or not. Other inputs may be seen by other devices such as the binocular (V\_bino), telescope (V\_tele), Charged Coupled Devices (CCD) camera (V\_ccd). Several ML techniques will be demonstrated in this work, especially the ones that we call static ML systems that do not require deep processing of data or temporal inputs [[2\]](#page-21-1). Some of the models that are used in this paper are: Support Vector Machine (SVM), Decision Tree, AdaBoost, Random Forest, logistic regression, Naïve Bayes, KNN, and Extra Trees. As in most ML problems, the frst stage will be feature selection. The training will be applied for the whole 12 months for the prediction of visibility of the new crescent Moon. The training will also be repeated based on classified/clustered continents and regions. Therefore, we may have a predictor for every region rather than having a single predictor for all regions in the dataset. This is an attempt to increase the accuracy by using the same predictor for geographically close reasons that are expected to have the new crescent Moon seen at the same time [\[3\]](#page-21-2). Moreover, in order to improve the accuracy, ensembled and hybrid techniques will be used. Meta heuristic optimization such as the Genetic Algorithm (GA) and the Particle Swarm Intelligence (PSO) are used to optimize the hyper parameters of the best performing ML algorithms [[4\]](#page-21-3). Meta heuristic optimization such as the Genetic Algorithm (GA) and the Particle Swarm Intelligence (PSO) are used to optimize the hyper parameters of the best performing ML algorithms. GA have proven successful in tackling numerous optimization challenges. PSO, on the other hand is known for its simplicity and efficiency in continuous optimization tasks. Both algorithms have shown high performance across a wide range of problems  $[4–7]$  $[4–7]$  $[4–7]$ . The idea is to, at least, reach fast convergence to the optimum hyperparameters if not much improvement has been achieved in the accuracy. This convergence to the optimum parameters is expected to be faster than the exhaustive grid search method.

Detecting the crescent using our proposed ML methods will not replace the traditional Islamic method of using the naked eye to "see" the new crescent Moon. For religious occasions such as the Holy month of Ramadan and the Eids, the new crescent Moon has to be seen by the naked eye. The ML outcome will definitely support any claims that the new crescent Moon has been seen. The sole target used in training the ML algorithms is  $V$ <sub>eye</sub> target (i.e., 0 not seen by eye or 1 seen by eye). Those values are usually captured from previous experiences. Therefore, the datasets on those ML methods have high credibility. Therefore, the proposed methods will train the predictors to make decisions based on Islamic law (seen by the eye). In Islamic events such as Ramadan or Eid the crescent must be seen by the naked-eye.

It is worth mentioning that investigations of minimizing the imbalance in the dataset will be conducted using upsampling and downsampling methods  $[8]$  $[8]$ . The fact that we treat the problem as a classifcation problem requires that we should try to cure the imbalance that we have in the original dataset. The imbalanced dataset also implies that the most appropriate performance metrics would be the F1 score, the AUC [Area Under the Receiver Operating Characteristic (ROC) Curve], the precision and the recall. It is not recommended to rely on accuracy metrics for this type of problem. Many studies emphasized the use of AUC for skewed data [[9\]](#page-21-6). Several previous works for predicting the start of lunar months are available in literature. Some authors used classical methods that depend on analytical and numerical calculations [\[10](#page-21-7)]. Others used simple supervised ML methods to gain high precision and recall values [\[11](#page-21-8)]. Al-Rajab et al. [[12](#page-21-9)] tried several ML methods on their own datasets and obtained encouraging results. Some works used image processing techniques and images caught using telescopes and binoculars [\[13](#page-21-10)]. Allawi [\[14](#page-21-11)] used a pattern recognizer method to estimate the presence of the new crescent Moon. In this work however, the goal is to achieve the following objectives:

- a) Improve the accuracy of ML techniques for the prediction of the new crescent Moon, investigating a wide spectrum of methods that vary in methodology and approach.
- b) Enhancing the existing methods by ensembling diferent models and hybridizing the existing ones with hyperparameters optimization.
- c) Dividing the datasets into clusters depending on geography or closeness of the countries.
- d) Using the (V\_tele), (V\_bino), and (V\_ccd) as possible inputs.

It should be noticed here that the target is V\_eye values and not any other detection values such as seen by the binocular or seen by CCD camera, or telescope.

The paper is divided into the following sections; ["Background and literature review"](#page-3-0) section is background and literature review, "[Dataset"](#page-5-0) section covers the dataset description, ["Methodology and experimentations](#page-6-0)" section presents the methodology and experimentation, "[Results and discussions"](#page-13-0) section presents the results and discussions, and fnally ["Conclusion and future work"](#page-20-0) section is the conclusion and future work.

#### <span id="page-3-0"></span>**Background and literature review**

There exist several calendar systems used by different civilizations such as Hijri Calendar, Solar Calendar, Gregorian Calendar. For instance, the solar calendar was restricted only to some Asian countries. Gregorian calendar, which is used worldwide [\[15](#page-21-12)]. However, Hijri calendar [\[16](#page-21-13)] represents an important reference to millions of people, Muslim, and Islamic states, across the world to identify signifcant religious events. In the Hijri calendar, a year is determined by 12 lunar months, each of which consists of approximately 29.530 days, accommodating the lunar phases. Consequently, the Hijri year encompasses approximately 354.360 days. This lunar-based system deviates slightly from the solar year in the Gregorian calendar, which spans around 365.25 days. Of these months, Muhrram, Rajab, Ramadan, Shawwāl, and Dhū al-Ḥijjah, which bear a unique signifcance for Muslims.

Before we start covering all research papers which tackled new Moon visibility, let us frst introduce the technical terms related to this research. Table [1](#page-3-1) summarizes all these terms, the formulas used to compute some features such as Age\_of\_Moon, Sun\_Moon\_ Lag, etc. can be found in [[12\]](#page-21-9).

The commencement of each Hijri month hinges on the sighting of the new crescent Moon. The Moon's position relative to the Earth and the Sun causes it to illuminate partially, a phenomenon known as Moon illumination, thereby altering its appearance or phase. Consequently, the Moon cycles through eight distinct phases, initiating from the

<b>Technical terms</b>	<b>Definition</b>
Age_of_Moon	Number of days (time elapsed) since the Moon was last in conjunction with the Sun
Altitude difference	Difference between the Moon altitude and the Sun altitude
Azimuth difference	Angular difference in azimuth between the Sun and the Moon. The azimuth of an object defines its position on the celestial sphere in the horizontal coordinate sys- tem, measured in degrees starting from 0 north and turning in a clockwise direction
Moon_conjunction	Alignment of the Moon between the Earth and the Sun, with its illuminated side facing away from us, making it invisible from Earth
Elongation	Angular distance between the Moon and the Sun
Moon_altitude	Height of the Moon above the horizon
Moon_azimuth	Compass discretion (north $= 0^\circ$ ) to the Moon's position in the sky
Sun altitude	Height of the Sun above the horizon
Sun azimuth	Compass direction (north $= 0^\circ$ ) to the Sun's position in the sky
Sun_Moon_Lag	Time difference between the positions of the Sun and the Moon as seen from Earth
V bino	New Moon was visible through binoculars
V_ccd	New Moon was visible through charge-coupled devices camera
V_eye	New Moon was visible by the eye
V tele	New Moon was visible through telescope

<span id="page-3-1"></span>**Table 1** Essential vocabulary for this research

unseen new Moon, progressing to the slender crescent illuminated by Sunlight, referred to as the new crescent Moon. This is depicted in the initial illustration of Fig. [1](#page-4-0). The Moon proceeds through subsequent phases until reaching its fnal phase, initiating a new cycle. Figure [1](#page-4-0), sourced from NASA's website, provides a visual representation of these phases [\[17](#page-21-14)].

Numerous studies have addressed the proposal to unify the Hijri calendar worldwide. For instance, lately, Mufd et al. [[18\]](#page-21-15) engaged in discussions concerning the unifcation of the Hijri calendar for the Southeast Asian region, specifcally focusing on Brunei, Indonesia, Malaysia, and Singapore. Wahidin [\[19](#page-22-0)] presented a literature review that integrates insights from several perspectives: hadith, astronomy, and sociology to investigate potential approaches for the unification of the Hijri calendar. The findings suggest that taking a comprehensive approach could accelerate the process of the Hijri calendar unifcation.

On the other hand, Maskufa et al. [\[20](#page-22-1)] found that the signifcant challenges when it comes to unifying the Hijri calendar, primarily stemming from a combination of fqh (Islamic jurisprudence) and astronomical factors. Hafez [\[21\]](#page-22-2) proposed a mathematical model for the Moon sighting based on the motion of the Moon relative to the position of the observer on the Earth. The model was mainly based on the calculation of the ratio of the thickness of the luminous part of the Moon to its diameter. The author validated the model against data given in the Um-Alqura calendar, Al-Tawqifat Alelhamih book and showed high accuracy of the results.

Another call to the use of science in determining the start of the Hijri month was from Hasan [\[22\]](#page-22-3) who presented a review highlighting the collaboration between Fiqh and empirical science in determining the start of the Hijri month in Indonesia, emphasizing the role of science in translating fqh concepts into practical guidelines, the interplay between fqh and scientific Moon sighting criteria, and fiqh validation of scientific efforts through testimonies. While it focuses on the years from 2010 to 2023, it provides valuable insights into methods and criteria development by the Hisab and Rukyat, with the potential for further

<span id="page-4-0"></span>

**Fig. 1** The different Moon phases [[17\]](#page-21-14)

research to deepen the understanding of this integration and its contribution to advancements in fqh and empirical science in Indonesia. In his study in 2018, Maskufa [[23](#page-22-4)] placed specific emphasis on the figh dimension, with a particular focus on astronomical figh. This aspect of fqh pertains to the regulation to determine when religious activities should take place based on the movements of the Moon and Sun.

Further researchers addressed the visibility of the new crescent Moon by employing image processing methods with the emergence of advanced telescope technologies and satellites such as Kepler, has expanded the possibilities for automating the analysis of observations, prompting the necessity to train computer systems to perform these tasks [\[24,](#page-22-5) [25](#page-22-6)]. For instance, Moshayedi et al. [\[26\]](#page-22-7) suggested a machine learning prototype system that employs a collection of Moon images to recognize and match the respective lunar phase represented in each image. Fakhar et al. [\[13](#page-21-10)] proposed an image processing-based method involving stages like noise elimination with a Gaussian flter, enhancing image quality around the crescent Moon area, and utilizing the circular Hough transform to extract the crescent's features. Sejzei and Jamzad [[27](#page-22-8)] created a Matlab toolbox designed to improve the visibility of a crescent Moon in images, aiding observers in its detection. Utama et al. [[28](#page-22-9)] utilized computer vision algorithms on video data to quickly and accurately identify the young lunar crescent's appearance. The process involved Gaussian Blur, Adaptive Thresholding for capturing the lunar crescent, and the Circular Hough Transform along with the OpenCV package for image extraction, processing, and detection.

Although, several studies have been conducted using AI techniques including ML and DL in several fields  $[29-33]$  $[29-33]$  $[29-33]$  and in particular the field of astronomy  $[34-36]$  $[34-36]$  $[34-36]$  but very few to almost no research focused on new crescent Moon prediction. For instance, Tafseer [[11](#page-21-8)] addressed the new crescent Moon visibility problem by employing machine learning algorithms, namely logistic regression (LR), Neural Network (NN), Support Vector Machine (SVM), and Random Forest (RF). In his study, he utilized a dataset collected from the Islamic Crescents Observation Project (ICOP) website [[37](#page-22-14)]. The dataset was augmented with additional features such as the age of the Moon, the Moon's lag time, the altitude difference, DAZ, the Moon phase, and atmospheric conditions. His results revealed that RF achieved the highest precision of 0.88 and a recall of 0.87. On the other hand, Allawi [[14](#page-21-11)], utilized a single ML algorithm, the Artifcial Neural Network (ANN) to predict the visibility of the new crescent Moon but confined to one particular region namely, Iraq. The algorithm yielded results with an accuracy level not less than 77%.

In our earlier study [[12](#page-21-9)], we conducted research to predict the start of the Ramadan month, employing classifcation supervised ML algorithms on a dataset acquired from ICOP and augmented with additional observations and features. Our results revealed that SVM, NN, and RF yielded promising prediction outcomes with an accuracy rate of 91%. In this research, we aim to extend our investigation to establish the Lunar Calendar for the entire Hijri Year of Islamic months, utilizing artifcial intelligence.

### <span id="page-5-0"></span>**Dataset**

To begin our research, we started by collecting and building our dataset. We ensured its comprehensiveness and adherence to necessary standards by sourcing raw data from the ICOP, specifcally from the "Crescent Observation Results". ICOP is a reliable source for lunar observations, recording data from numerous cities and countries worldwide. Their extensive network of observers guarantees accurate documentation of new crescent Moon sightings, providing a diverse range of observations from diferent geographical regions.

We manually built our own dataset from those observations, focusing on various features as outlined in our previous study  $[12]$  $[12]$ . These features include Hijri Day, Conjunction Time, Date, Country, City, State, Atmosphere, V\_eye, V\_bino, V\_tele, V\_ccd, Longitude, Latitude, Sunset, Moonset, Sun\_Moon\_Lag, Age\_of\_the\_Moon, Moon\_altitude, Sun altitude, Altitude difference, Moon azimuth, Sun azimuth, Azimuth difference, Elongation, and Illumination. The formulas for calculating features such as Age\_of\_the\_Moon, Sun\_Moon\_Lag, Altitude\_diference, Azimuth\_diference, and Illumination can also be found in [\[12](#page-21-9)].

We processed and translated data such as country and city names from Arabic to English, standardizing spelling variations and handling abbreviations as necessary and explained in the data preparation "[Dataset preparation"](#page-7-0) section. In our previous study [[12\]](#page-21-9), we focused exclusively on data for the month of Ramadan. In this work, we have signifcantly expanded that initial dataset to now include comprehensive data from all months over a 13-year period, from 2010 (1431 Hijri) to 2023 (1444 Hijri).

When deciding which features to include, we used commonly accepted data for this type of study [[38\]](#page-22-15) and augmented it using the Accurate Times 5.6 software [[37\]](#page-22-14). By inputting the country, city, and date, the output included extra important features such as sunset, moonset, altitude, and longitude. Our fnal dataset comprises a total of 2085 observations from 47 diferent countries, as shown in Fig. [2](#page-6-1). A sample of the dataset is presented in Fig. [3.](#page-7-1)

#### <span id="page-6-0"></span>**Methodology and experimentations**

In this research paper, our main contribution is focused on achieving the precision of the new crescent Moon prediction through an exhaustive exploration of diverse machine learning techniques and approaches. Additionally, based on the obtained results, we attempted to combine diferent models and optimize hyperparameters to achieve better prediction results. Furthermore, we intended to segment datasets into clusters based on geographical or country proximity factors.



<span id="page-6-1"></span>**Fig. 2** Map of countries included in Moon observation dataset

index hijn year	Month	conjunction	Conjunction Time	date	hijri date	country	city	state	atmosphere v_eye v_bino v_tele v_ccd				longitude	latitude	sunset	moonset	Sun Moon Lag	Age of Moon
	31 Dhul Hijah	06/11/2010 4:52	452	06/11/2010	29/11/1431 Algeria		Ghardaia	Ghardaia				o	48000	23 00 0	1834	1920	0.76666667	137
		31 Dhul Hijah 06/11/2010 4:52	4.52		06/11/2010 29/11/1431 Australia		Sydney	New South Wales	0.85			$\theta$	0.11,00.0	$-3357000$	122	847	1.416666667	25
		31 Dhul Hijah 06/11/2010 4:52	452		06/11/2010 29/11/1431 Bancladesh		Dhaka	Dhaka				$\theta$	0.20.16.0	46.10.0	1236	13:22	0.766666667	7.733333333
		31 Dhul Hijah 06/11/2010 4:52	452		06/11/2010 29/11/1431 Ghana		Accra	Greater Region	0.35		0	$^{\circ}$	0.10000	36 00 0	18:15 19:31		1,26666667	13 38333333
		31 Dhul Hijah 06/11/2010 4:52	452		06/11/2010 29/11/1431 Iran		Esfahan	Esfahan	0.35				0.40.45.0	39.05.0	15:30	16.11	0.683333333	10.63333333
		31 Dhul Hijah 06/11/2010 4:52	452		06/11/2010 29/11/1431 Irao		Moosel	Ninavah	0.85		$^{\circ}$	$\theta$	0 08:01.0	20:16.0	16:06 16:42		0.6	11,2333333
		31 Dhul Hijah 06/11/2010 4:52	452		06/11/2010 29/11/1431 Oman		Muscat	Muscat	0.35	$\mathbf{r}$			3536.0	36.48.0	14:42	15:33	0.85	9833333333
		31 Dhul Hijah 06/11/2010 4:52	452			06/11/2010 29/11/1431 Saudi Arabia	<b>Jahal Khandamal</b>	Makkah	0.85			$\theta$	0 38 59 0	59530	13.55	13.55	00	905
		32 Dhul Hijah 26/10/2011 19:56	19.56		26/10/2011 28/11/1432	Tanzania	Daressalaam	Tanzania	0.35				0.16110	$-0649230$	15:16	14:57	$-0.316666667$	-4 666666667
		32 Dhul Hijah 26/10/2011 19:56 19:56				26/10/2011 28/11/1432 United States	Ganesville	Florida				0	0.8217000	41:00.0	22.48	2235	$-0.216666667$	286666667

<span id="page-7-1"></span>**Fig. 3** Sample of the observation features included in our dataset



<span id="page-7-2"></span>**Fig. 4** The proposed framework model

Our overall system architecture is presented in Fig. [4.](#page-7-2) As shown in the fgure, the experiment design goes over diferent steps, starting from the dataset collection and building, data pre-processing and feature selection, training and validating the data through a combination of ML models and approaches (as it will be explained later).

## <span id="page-7-0"></span>**Dataset preparation**

The data we collected went through a pre-processing phase to guarantee its appropriateness for analysis. Prior to conducting our experiments, we took several pre-processing steps:

- Missing values: all missing or null values were eliminated to ensure data completeness.
- Invalid values: rows with negative Sun\_Moon\_Lag values were eliminated, as these are not valid for new crescent Moon determination.
- Data validity: rows with a Hijri Date of 28 were eliminated (although the observation of the new crescent Moon was recorded on ICOP, including the 28th of Hijri months, it can only be determined on the 29th or 30th of a Hijri month.)
- Data standardization and normalization: we standardized the data using min–max standard scaler normalization, scaling it to values between 0 and 1. This normalization technique is commonly used to ensure that all data is on a consistent scale, making it easier to compare diferent variables during analysis.
- We translated countries from Arabic to English and standardized spellings to match those used by the Accurate Times 5.6 software (e.g., Mecca to Makkah).
- Abbreviations were expanded to their full names.
- The atmosphere feature was converted to numeric values (Superb = 1, clear = 0.85, hazy =  $0.6$ , Very hazy =  $0.5$ ).
- The observation feature, V\_eye, was converted from categorical to numeric to enable better understanding, ftting, evaluation, and extraction of valuable information by the machine learning model.
- Total Number of observations after data processing reduced to 1779 sample points.

To decide on the performance metrics to consider in our study, we had to analyze our data. We are interested in predicting the visibility of the Moon using the naked eye, for that we chose the V\_eye feature as our main prediction label. Plotting the V\_eye values from the 1779 observations, as depicted in Fig. [5,](#page-8-0) it was found that 68.4%, with 1216 instances out of a total of 1779 of the observations were seen by the naked-eye (V eye = 1), and 31.6% with 563 instances out of a total of 1779 of the same observations were unseen (V  $eye=0$ ). Consequently, to remedy the dataset unbalance, we used the F1-score and precision for the model performance evaluations.

## **Feature selection**

After data balancing exploration, feature selection algorithms were explored to increase the possibility of reducing the complexity of the problem by reducing the number of parameters fed in. Diferent feature selection algorithms were applied [[39\]](#page-22-16) where some focus on linear and other focuses on the non-linear characteristics of the data. Each feature selection method was employed with a range of features to identify the optimal subset for training the models. These subsets were then used to train various machine learning algorithms, ensuring that each model had access to the most relevant features for making predictions. A brief description of these feature selection algorithms used in our experiments is provided below.

• *The ANOVA F-test* this method is used to evaluate the significance of each parameter in explaining the variation of a target variable. It is calculated by taking the ratio of



<span id="page-8-0"></span>**Fig. 5** The distribution of samples across the two classes (V\_eye = 0 and V\_eye = 1) in the dataset

variance between the parameters to the variance within each parameter. The result of the ratio (F1 score) is then compared with the P-value to assess the statistical signifcance of their relationship with the target value, the higher the f-test score the stronger the dependency.

- *The mutual information* this method can capture both linear and non-linear relationships between the features and the target, without making assumptions regarding the distribution. This method relies on the entropy formula to check and evaluate the relationship between the features and target variable. The selected features are based on the higher mutual information score.
- *Te recursive feature elimination with logistic regression (RFE)* method combines the principle of back propagation and logistic regression by removing the least important features until it reaches the set number of features. The corresponding coefficient is the controlling metric which arranges the features based on importance.
- *The least absolute shrinkage and selection operator (LASSO)* is used for feature selection and regularization by identifying the most important feature based on a modified linear regression cost function. The cost function assigns zero weights on the less important feature within the iterations.
- *The principal component analysis* is a dimensionality reduction technique in which it transforms high dimensional data to lower dimensional representations. The dimensionality reduction is achieved by fnding the linear combinations of the features which would capture the maximum variance.
- The correlation coefficient is used to assess the degree of linear association between the target variables and individual features. It is also used to assess the degree of linear association between the features between themselves. The correlation coefficient is calculated and is usually represented as a correlation matrix. It works well when the features are associated with a linear relationship.
- *The gradient boosting importance* it works by applying an ensemble of decision trees, in which the features are split based on the value. The features that are frequently used for splitting the tree are considered of more importance. However, this method works on the predictive power of features not its causation. It can handle class imbalances which is the case in this problem.
- *The linear support vector classifier's* importance aims at finding the best hyperplane which accurately separates the data points that belong to diferent classes. During the training process each feature is assigned a certain weight which indicates the contribution of each parameter.
- The random forest importance classifier create an ensemble machine learning algorithm which can be used for feature selection. It calculates feature importance based on the importance score of each feature in reducing the error of the decision tree. It works best with numerical and categorical features.
- The extra trees importance is another ensemble method in which it helps in feature selection based on their importance. It applies further randomness in the tree building process.

Each feature selection method was employed with a range of features to identify the optimal subset for training the models. Tese subsets were then used to train various machine learning algorithms, ensuring that each model had access to the most relevant features for making predictions.

### **Implementation details**

Our models were implemented and trained using the Collaboratory notebook (Colab), a cloud-based tool provided by Google Research, ofering a lightweight and free envi-ronment [\[40\]](#page-22-17). The programming language used was Python, which incorporates several Python libraries such as Pandas, NumPy, Scikit-learn, Matplotlib, Seaborn, among others. Pandas facilitated data manipulation and analysis, while NumPy handled computations. Matplotlib and Seaborn were utilized for data visualization purposes. For the implementation of machine learning algorithms, we relied on the Scikit-learn library.

## <span id="page-10-0"></span>**Experimental setup**

For our experiments, we employed commonly used machine learning algorithms for classification purposes. These included Naïve Bayes (NB), Support Vector Machine (SVM), Random Forest (RF), gradient boosting (GB), logistic regression (LR), K-Nearest Neighbor (KNN), Decision Tree (DT), Xgboost, AdaBoost, and Extra Trees. In all our experiments we used tenfold cross validation. The latter exhibits important benefts which include model performance assessment across diferent subsets of the data, reduction of overftting risk, provision of robustness against data random variations, and is a widely adopted approach in the feld. In the initial phase of our experiments, we conducted a correlation analysis among the features in our dataset. The results were visualized through a correlation heatmap as illustrated in Fig. [6,](#page-11-0) where important features exhibit high correlation. Particularly, the features such as "illumination" shows signifcant correlations with "Sun\_Moon\_lag", "Moon\_altitude", "Altitude\_difference", "elongation", and "Azimuth\_diference". Tis indicates a strong linear relationship between these features and the "illumination". Notably, the most signifcant correlation is detected between "Sun\_Moon\_lag" and both "Moon\_altitude" and "Altitude\_diference", each having a correlation coefficient of 0.98. It is worth noting that both "Sunset hour" and "Moonset\_hour" are highly dependent on "Age\_of\_Moon" (0.54, 0.55) and "longitude" (0.7, 0.72). Given the results obtained from the heatmap, this suggests the use of the non-linear ML models which are found more suitable for our analysis.

In our research, we conducted a series of experiments aimed at comprehensively analyzing and optimizing our model's performance. Experiment 1 examined how excluding certain features afected our predictions. Experiment 2 included all features to provide a complete interpretation for comparison. Experiment 3 utilized feature selection and grid search models to enhance the model's performance by relying on relevant features and fne-tuning hyperparameters. Experiment 4 explored the benefts of ensemble learning, combining multiple models for improved accuracy and robustness. Experiment 5 employed a hybridized approach to efficiently optimize hyperparameters. Experiment 6 investigated a region-based approach to capture geographically specific patterns. These experiments contributed to the iterative refnement and enhancement of our predictive model, advancing our understanding and applicability in the target domain.



<span id="page-11-0"></span>**Fig. 6** Dataset correlation Heatmap

## *Experiment 1: analysis of full 12‑month dataset excluding V\_ccd, V\_tele, and V\_bino features*

In Experiment 1, our approach involved the application of the previously mentioned ML algorithms ("[Experimental setup"](#page-10-0) section). This analysis was conducted using the default hyperparameters for each of the ML algorithms. It is important to note that the set of input parameters in this experiment excluded three specifc features: V\_tele, V\_bino, and V\_ccd. Tis exclusion was aimed at assessing the predictive capabilities of these models under modifed dataset conditions.

#### *Experiment 2: analysis of full 12‑month dataset with all feature included*

In this experiment, we chose to include all the available features of the dataset. Tis approach was undertaken to gain more comprehensive insights into the dataset characteristics. With all the features present, our objective was to assess whether the overall performance and predictive capabilities of the models improved. The same 10 ML algorithms were applied.

### *Experiment 3: feature selection and model optimization using grid search*

In this experiment, the dataset underwent a systematic process involving feature selection and grid search of various ML models which are presented in Fig. [7](#page-12-0). Tis process aimed to identify the most efective model in conjunction with the optimal set of features. The experimental procedure started by defining a number of features,  $x$ , to be selected among the 24 features in our dataset (full dataset). In our case *x* was set as 7. Within each iteration of this process, the feature selection algorithm selected a specifc number of  $x$  features. These selected features were then input into the grid search technique to fne-tune the hyperparameters of a model.

## *Experiment 4: investigating enhanced predictive performance through ensemble modeling*

In this experiment, we aimed to investigate the potential enhancement of predictive results through ensemble technique. This approach integrated the four best-performing models based on the top F1 scores released from experiment 3, each with their respective input features, and passed them through a soft voting classifer, which combined the average of the votes. The objective was to assess whether this ensemble methodology could further augment the performance outcomes. To evaluate the value of the ensemble technique, we benchmarked its results against those of the top 4 performing models obtained in the previous experiment 3. Tis comparative analysis provided insights into the viability and efectiveness of employing an ensemble of models for improved predictive capabilities in our dataset.

## *Experiment 5: exploring hybrid optimization techniques for enhancing ml model performance*

Tis experimental setup aimed to explore the potential enhancement of the top-performing model, from the results obtained in experiment 3 through hyperparameter optimization using Genetic Algorithm (GA) and Particle Swarm Optimization (PSO) techniques. These optimization methods offer advantages in fine-tuning model parameters to maximize performance, thereby helping in a more refned and efective model. To maintain consistency in the experimental conditions, the same set of 14 features, as previously identifed, will be utilized for this phase of the study.



<span id="page-12-0"></span>**Fig. 7** Feature selection and grid search pipeline



<span id="page-13-1"></span>



<span id="page-13-2"></span>**Fig. 8** Map of countries segmented into regions

### *Experiment 6: investigating the infuence of geographical factors: a region‑based approach*

In this experiment, we explored deeper the infuence of geographical features such as longitude and latitude on our models. This exploration was driven by the observation in Experiment 3. We categorized the observations, derived from 47 countries, into fve distinct regions: Middle East, Africa, Asia, Europe, and Other. The region segmentation is detailed in Table [2](#page-13-1) and illustrated in Fig. [8](#page-13-2).

#### <span id="page-13-0"></span>**Results and discussions**

In all our experiments, we used the F1 score and the mean of the area under the curve (AUC) as the main evaluation performance metrics to assess our prediction results.

For experiment 1, the performance of the ML models considered in the analysis is illustrated in Table  $3$ . The GB outperforms all the other models, boasting the highest F1 score of 0.824845 and an AUC of 0.764184, indicating its ability to handle the complexity of the dataset. On the other hand, NB exhibits the lowest performance with an F1 score of 0.608916 and an AUC of 0.617499, showing its struggle with the dataset complexity. In general, ensemble methods like GB, RF, and XGBoost demonstrate robust performance, indicating complex patterns within the data. SVM also shows strong performance with a high F1 score, indicating its capability to manage high-dimensional data and accurately detect true positives.

Experiment 2 involved the comprehensive analysis of the full dataset, inclusive of features V\_ccd, V\_bino, and V\_tele, resulting in a noticeable enhancement in the performance metrics across all applied ML models, as illustrated in Table [4](#page-14-1). Tis

<b>ML</b> algorithm	Mean F1 score	<b>Mean AUC</b>
<b>NB</b>	0.608916	0.617499
<b>SVM</b>	0.820968	0.706106
<b>RF</b>	0.808872	0.719215
GB	0.824845	0.764184
<b>LR</b>	0.800651	0.637309
<b>KNN</b>	0.781338	0.676872
DT	0.740491	0.610156
XGBoost	0.787297	0.717707
AdaBoost	0.814294	0.750441
Extra trees	0.800238	0.683953

<span id="page-14-0"></span>**Table 3** Results of experiment 1 (The bold values in the table indicate the highest performance metrics (Mean F1 Score and Mean AUC) achieved by the machine learning models in experiment 1)

<span id="page-14-1"></span>**Table 4** Results of experiment 2 (The bold values in the table indicate the highest performance metrics (Mean F1 Score and Mean AUC) achieved by the machine learning model in experiment 2)

Algorithm	Mean F1 score	<b>Mean AUC</b>
<b>NB</b>	0.771363	0.798937
<b>SVM</b>	0.876087	0.874549
<b>RF</b>	0.870211	0.864888
GB	0.882469	0.901009
LR	0.858785	0.833284
<b>KNN</b>	0.866233	0.857882
DT	0.837237	0.752882
XGBoost	0.854933	0.882258
AdaBoost	0.846337	0.807547
<b>Extra trees</b>	0.87041	0.868849

improvement distinctly indicated the signifcant role these features played in augmenting the predictive F1 score. In particular, GB emerged as the most efective model, achieving an F1 score of 0.882469 and an AUC of 0.901009. Tis better performance indicated the compatibility of the dataset with sophisticated ensemble techniques, as also evidenced by the high efficiency of models like SVM, RF, and XGBoost. However, the NB, despite its improvement to an F1 score of 0.771363 and AUC of 0.798937, remained the least efective among its counterparts.

Experiment 3 records the outcomes of each iteration, collecting a comprehensive result from the dataset. Upon completion of all iterations, which cover more than 7000 combinations of selected features and model configurations. The most effective four combinations, based on predetermined performance metrics are shown in Table [5](#page-15-0). Tis method enables a thorough examination of the models' efectiveness across diverse feature sets. It also enhances the comprehension of feature signifcance within the dataset.

Table [5](#page-15-0) specifcally describes the performance metrics—Mean F1 Score and Mean AUC of the four most performant models which are, the Extra Trees algorithm and the SVM.

The analysis of the dataset, employing the different feature selection techniques, has improved the predictive performance across several ML models. This enhancement indicates the importance of specific features, such as "V\_bino", "V\_tele", "V\_ccd", "longitude", and "latitude", among others, in improving the overall model efficiency as depicted by their high respective F1 scores.

It has been found that the Extra Trees model, optimized with parameters of max\_ depth: 10 and n\_estimators: 50, and using a set of 14 features selected through ANOVA F-test, stands out with an impressive mean F1 score of 0.887872, and a mean AUC of 0.906242. The marginal differences in  $F1$  scores across models, being less than 0.02%, illustrate the impact of feature selection on model accuracy. On the other hand, the SVM model, employing the Random Forest Importance for feature selection and optimized with parameters of C: 1 and kernel: 'rbf', also shows promising results. When applied with 10 selected features, it achieves a comparable mean F1 score of 0.887848, and a mean AUC of 0.888855. Notably, even with a reduced feature set of 8 or 9 features, the SVM model maintains high performance levels.

These results underscore the effectiveness of employing targeted feature selection methods and fine-tuning model parameters. The effectiveness of ensemble methods and the SVM's kernel trick suggest that the dataset likely exhibits non-linear relationships and interactions among its features, which are more accurately captured by models capable of handling high-dimensional spaces and complex feature interactions. These results highlight how useful it is to carefully choose specifc features and adjust model parameters. The success of combining multiple models and using advanced techniques indicates that the dataset may contain complex relationships between its features. Models that can



<span id="page-15-0"></span>**Table 5** Results of feature selection models and grid search

#### <span id="page-16-0"></span>**Table 6** Ensemble model



#### <span id="page-16-1"></span>**Table 7** Comparison with hybrid models



process lots of diferent features and their interactions tend to perform better as it was illustrated in our case.

In experiment 4, we extended our analysis to include ensemble models. To evaluate the importance of this technique, we will benchmark its results against those obtained from experiments 3 (Table [5\)](#page-15-0). Tis comparative analysis aims to refect on how using a combination of models can enhance the predictive capabilities within our dataset. It is illustrated from Table [6](#page-16-0), that the Ensemble Model demonstrates the highest F1 score of 0.888058 coupled with an AUC of 0.907482, compared to the results of (Table [5\)](#page-15-0).

The objective of experiment 5, is to identify the most optimal hyperparameter combination that maximizes model performance. A defned range of hyperparameter values for the Extra trees model, specifcally 'n\_estimators' within the range of 10 to 200 and 'max\_depth' within the range of 1 to 50, will be subjected to both GA and PSO optimization processes. Tis comparative study aims to determine if the GA and PSO-based methods for hyperparameter optimization can outperform the performance standards established by the grid search technique. The outcomes of this optimization will then be compared with the results obtained from Experiment 3, as outlined in Table [7.](#page-16-1)

The results depicted in Table [7](#page-16-1) clarify a crucial aspect of our ML model optimization. The convergence patterns observed from both the GA and PSO optimization indicate that the performance level stays about the same without much improvement  $[41]$  $[41]$  $[41]$ . However, the convergence behavior within 40 generations by the GA and the PSO shows the efficiency of those algorithms in reaching the optimal solution without the need for the exhaustive grid search that is usually used to fnd the optimum hyperparameters values. Figures [9](#page-17-0) and [10](#page-17-1) depict the convergence behavior of GA and PSO. Since we are applying a tenfold cross validation it would be advisable to use efficient techniques such as the GA and the PSO for parameters hyper tuning optimization. The GA and the PSO were applied on the best performer algorithm which is the "Model 1/Extra trees". The convergence to the optimum hyperparameters was much faster as compared to the exhaustive grid search. Both algorithms used a cost function that is based on the F1 and AUC values. Tis work leads to propose a possible direction for future work, a more expansive approach to hyperparameter tuning across the spectrum of the 10 ML models employed, this will help us investigate deeply the hyper parameter space, potentially uncovering more efective model confgurations.



<span id="page-17-0"></span>**Fig. 9** Convergence behavior of the GA over the generations; average ftness (to the left) and standard deviation (to the right)



<span id="page-17-1"></span>**Fig. 10** Convergence behavior of the PSO algorithm over the generations

<span id="page-17-2"></span>



In experiment 6, we divided our dataset into regions. The distribution of  $V$  eye labels within these regions is presented in Table [8](#page-17-2). It is worth noting, we can see that there is a relative balance of datasets in the Middle East and Europe, in contrast to the signifcant imbalance in the rest of the regions.

To assess the impact of the regional segmentation on the predictive performance, we applied the same methodology adopted in experiment 3. The top performing models for each region are illustrated in Table [9](#page-18-0).

The analysis of the regionalized data, as reflected in Table [9](#page-18-0), shows an improvement in the F1 score for some regions such as Africa and Asia, and this compared to the results obtained from experiment 3. However, there is a decline in the AUC scores. This can be due to a reduced number of samples, which can be attributed to the limited false label ('V\_eye = 0') occurrences in the datasets. When comparing the performance metrics of individual regions to the 'All Regions' model, it is visible that <span id="page-18-0"></span>**Table 9** Model evaluation of different regions (The bold and italic values represent the highest Mean F1 Score and Mean AUC, respectively, within each region's model evaluation. These values highlight the best-performing model for that specifc region, demonstrating the model's ability to most accurately predict the visibility of the new crescent Moon based on the region-specifc dataset)



while some regions beneft from the regionalization approach in terms of F1 Score, the AUC scores are reduced. This observation indicates the impact of the dataset size on model performance. The number of samples in the different regions was mostly unbalanced.

Figure [11](#page-19-0) depicts a comprehensive view of all the scores via the diferent ML models and regions. The RF, GB, KNN, and Xtree models were the top performers as proven by shown results and the previous results. Probably not much improvement has been noticed by allocating an independent predictor for each one of the 5 regions, however, based on comparisons with the all-regions predictor, some improvement has been noticed for Africa and the "other" region with regard to the F1 score when using the KNN model. It is worth mentioning that the number of samples were still unbalanced despite the region's division process. In fact, some regions had closely balanced samples such as the "Middle East" but still no improvement was noticed for that region. Tis leads us to reach some conclusions regarding the training data balance characteristic as will be mentioned in the conclusion.



<span id="page-19-0"></span>**Fig. 11** All score values as per region and ML models

This study aimed to establish the Hijri calendar for a span of 13 years, incorporating various countries from diverse regions worldwide. In contrast to our previous research in [\[12\]](#page-21-9), which solely focused on the month of Ramadan, this study encompasses all months of the Hijri year, utilizing an updated dataset spanning the entire 13-year period. To our knowledge, only two studies have been conducted to predict the visibility of the new crescent Moon. Allawi  $[14]$  $[14]$  $[14]$ , only one deep learning algorithm, Convolutional Neural Network (CNN), was employed, limited to a single country over a 4-year period. Tafseer  $[11]$  $[11]$ , four supervised learning algorithms were applied to the dataset. Our study represents the first comprehensive examination of all 12 Hijri months over a 13-year period from x to y. We conducted an extensive analysis employing a wide array of ML algorithms and techniques, including default parameters to grid search, model ensemble, model hyperparameter hybridization, then ending with regional splitting of the data with the main purpose of extracting the best performing model and maximizing its performance.

Overall, the GB and the Extra Trees outperformed the other machine learning models. It is apparent that decision-tree based models (GB and Extra Trees) are more appropriate for this type of problems and due to their ensemble nature, their ability to capture the interactions of complex features, and the ability to manage the imbalanced data effectively. Other machine learning models that have a connectionist nature such as NN (that use weight tuning approach) show less success as sometimes it is more difficult to find a surface that separates between the rough input regions. All these made the GB and the Extra Trees models well suited for predicting the visibility of the new crescent Moon.

The results show that the GB and Extra Trees models outperform the other models considered in this analysis. Each model processes data differently depending on the importance of specific features. In terms of feature selection, the Extra Trees model uses a broader set of features, which might explain why it performs slightly better than the GB model, which prioritizes features that reduce the overall error. The hyperparameter tuning using the GA and PSO metaheuristic algorithms helps the models learn faster and work more effectively, resulting in the superior performance of the Extra Trees model.

#### <span id="page-20-0"></span>**Conclusion and future work**

In this paper, we present ML-based framework that contributes signifcantly toward the realization of a globally unifed lunar calendar. Various ML algorithms were considered, and several experiments were conducted on the dataset, starting with default parameters, proceeding to Grid Search, Ensemble Modeling, model hyperparameter hybridization, and regional data segmentation. The primary objective was to identify and optimize the best-performing predictive model for the new crescent Moon visibility. Results revealed that the GB model outperforms the other models considered in this study in terms of F1 score and AUC. However, with feature selection identifed through the ANOVA F-test and optimized parameters, the Extra Trees model turns out to be the best-performing predictive model. The Ensemble model and hybridization, on the other hand, although they exhibit slight improvement compared to the other models, showed faster convergence to the optimum values of the models hyperparameters. Geographical segmentation of the dataset enhanced predictive performance in certain areas, such as Africa and Asia. A noteworthy point is that a balanced dataset is not sufficient to achieve improved performance in the segmented regions when compared to experiments that take all regions in the dataset into account.

For future work, we will consider incorporating new optimization techniques such as Diferential Evolution (DE) and Gradient-based Optimization algorithms. Additionally, data synthesis could be used to generate more valuable training and testing data. Collecting more data from diferent regions will be necessary to further explore the geographical segmentation of the dataset and train more robust ML models. Although the machine learning methods we have employed are advanced, DL models capable of capturing spatial and temporal variations in inputs could also be applied.

To extend our study, we plan to include real-time data for predictions using ongoing observations, particularly during the conjunction period. The input data can be a live stream or stored videos providing continuous observations, similar to real-life scenarios. While this method may require more processing storage and power, it offers a more comprehensive approach by utilizing rich spatial and temporal information to reach a decision.

#### **Abbreviations**





#### **Author contributions**

S.L.: contributed to the concept, methodology, analysis, and drafting of the paper. M.A.: contributed to the concept, methodology, analysis, drafting of the paper, development of the code, and supervised the research project. R.A.Z.: contributed to the concept, methodology, analysis, and drafting of the paper. M.R.: contributed to the development of the code used in the research and the validation of the results. S.L., M.A., and R.A.Z.: involved in data collection and interpretation.

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#### **Data availability**

The datasets used and/or analysed during the current study available from the corresponding author on reasonable request.

#### **Declarations**

#### **Competing interests**

The authors declare no competing interests.

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